Precision physics and antimatter

Summary – Day 1

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 Precision experiments are a complementary way of searching new physics





Summary – Day 1

- Comparing properties of conjugate particle-antiparticle pairs provide stringent tests of CPT symmetry
- Which process gave raise to the baryongenesis in the early universe?
- Search for CPT-violating interactions

$$(i\gamma^{\mu}\partial_{\mu} - a_{\mu}\gamma^{\mu} - b_{\mu}\gamma_{5}\gamma^{\mu} - m)\psi = 0$$

- Frequency difference in transitions involving spin transitions for fermion pairs
- Diurnal oscillations of observables



Some questions that were asked...

- Why should the CPT-odd fields be fixed in space?
 - Answer: It is an assumption by theory, which has to be experimentally tested.
 - It is a plausible assumption if you consider an analogy to the isotropic cosmic microwave background (CMB). CMB is a boson field (photons) which was released when the universe became transparent for photons (kT < 13.5 eV, hydrogen ionization energy). In analogy a "heavy boson field", making up the Standard Model Extension Field b, which is weakly interacting with matter could have been emitted in the early universe. The SME fields correspond in the quantum field theory picture to bosons which mediate an unknown CPT-odd force to fermions (electrons, protons, etc.)
 - The CMB looks like a black-body radiation with ~2.7 K temperature. It is isotropic with 10^-4 fluctuations
 reflecting temperature fluctuations in the thermal equilibrium in the early universe. If the "heavy boson field"
 reflects the same amount of fluctuations it is reasonable to assume the the SME fields have a rather constant
 value on the trajectory of the earth.
- If CP-violation is small, and if CPT-violation makes if it exists at all an even smaller contribution, how can it possibly explain the Baryon asymmetry?
 - Answer: Baryon excess generation via CP violation can only take place in thermal non-equilibrium, whereas baryon excess via CPT-violation can be permanently generated also in thermal equilibrium.
 - The contribution of this process can be hardly estimated without knowing the mechanism generating CPTviolation. The quantum field theories we know come with a wide range of coupling constants and masses, therefore a baryogenesis from CPT-violation based on the current measured constraints is not excluded.

Precision physics and antimatter

Part 2.1 Production and confinement of cold antiparticles

The Antiproton Decelerator of CERN



Physics in the AD/ELENA-facility



BASE, ATRAP, Fundamental properties of the antiproton

ALPHA, ATRAP, Spectroscopy of 1S-2S in antihydrogen

ASACUSA, ALPHA Spectroscopy of GS-HFS in antihydrogen

ASACUSA Antiprotonic helium spectroscopy

ALPHA, AEgIS, GBAR Test free fall/equivalence principle with antihydrogen

M. Hori, J. Walz, Prog. Part. Nucl. Phys. 72, 206-253 (2013).

Why do we need antiparticles at rest?

• Precision experiments:

"Never measure anything but frequency" – A. Schawlow

- Observation time ↔ Fourier limit Confinement increases the observation time
- Systematic limits often scale with temperature
 - Laser spectroscopy: Doppler shift
 - Single (anti)particle experiments: Trap imperfections, relativistic shift
- Antihydrogen synthesis: Recombination cross-section $\propto T^{-x}$, x > 0

Antiproton decelerator in 2015



Proton Synchrotron beam on irdium target: Pair production of antiprotons

Focusing and separation of antiprotons

Antiproton decelerator: Synchrotron for antiprotons

Decelerates from 3.6 GeV to 5.3 MeV by

1.) Rf-deceleration
 2.) Stochastic cooling
 3.) Electron cooling

Ejection to experiments: 30 million antiprotons in 200 ns every 110 s

Production of Antiprotons

Pair creation in HEP reaction:

$$p + p + E_{kin} = p + \bar{p} + p + p$$

Fixed target threshold energy for antiproton production

$$E_{kin} = 5.63 \text{ GeV}$$

Differential cross-section, maximum



Higher energy gives higher yield, but the antiproton momentum (~center of mass momentum) increases!

Design considerations



0.1

3.57 GeV/c

Collection at low energy not possible

D. Moehl, Hyp. Int. 109, 33 (1997).

10

Antiproton collection momentum (GeV/c)

The antiproton production target



Target: Highest density, high melting point, stability against thermal shock Limit in production yield: target heating by the proton beam (T < 1800 K)

The challenge: Cooling by ~10 orders of magnitude

Antiproton production: ~ 3.7 GeV

AD final energy: ~ 5.3 MeV

Trap injection energy: < 5 keV

Present cooling limit: 0.5 meV

rf-deceleration

Stochastic Cooling

Electron Cooling

Degrader

Sympathetic cooling Resistive cooling

Laser-cooled ions?

AD Cycle



- Antiprotons injected into the AD: finite phase, energy, momentum, position spread.
- Leads to betatron oscillation and instabilities
- Liouville's theorem: the phase space is conserved!
- Adiabatic deceleration further increases the transverse emittance
- Cooling mechanisms are needed!

Stochastic Cooling

Reduce betatron oscillations by signal pickup and active electronic feedback





AD Cycle



Electron Cooling

Electron in a strong magnetic field: Cyclotron Motion Permanently accelerated charge -> Maxwell equations -> electron irradiated photons

Typical magnetic fields (T) COOLDOWN to ambient temperatures within 0.1s

$$\frac{\mathrm{d}E}{\mathrm{d}t} = -\frac{e^2 a^2}{6\pi\epsilon_0 c^3} \qquad \qquad \tau = \frac{3\pi\epsilon_0 m^3 c^3}{e^4} \frac{1}{B^2}$$

Electron cooling



Note: Cooling limit

The antiprotons are only cooled on a small fraction of their flight path. On the remaining orbit heating, which is at low energies dominated by space-charge driven expansion of the antiproton bunch, counteracts the cooling process.

Electron cooling





AD electron cooler, photo: CERN Courier

electron gun
 electron collector
 central drift tube
 clearing electrodes

(5) gun solenoid
(6) expansion solenoid
(7) toroid
(8) cooling solenoid

(9) collector solenoid
 (0) sputter ion pumps
 Long cooling times
 (order of s) needed!

lons interact 10⁶ 1/s with a collinear beam of cold electrons.

Properties of the cold ions: momentum spread $\Delta p/p = 10^{-4} - 10^{-5}$ diameter d = 2 mm

AD Performance

QTY	#
Start Energy	3.5 GeV
Stop Energy	5.3 MeV
Efficiency	Typical 80%
Cycle length	120s
Bunch length	150ns
Particle per bunch	about 30.000.000
Catching Efficiency	< 1%

Some comments:

Slow AD cycles – one information in 2 minutes debugging is time consuming!

In one week, 3 experiments run in sequence for 8 hours shifts

Catching efficiency of 5.3 MeV antiprotons is low!



- Deceleration of antiprotons from 5.3 MeV to 100 keV to improve efficiency of experiments
- Circumference 30.4 m (1/6 the size of the AD), magnetic ring and electrostatic extraction lines
- Challenges related to low energy as field quality of magnets operated with very low fields

Courtesy of C. Carli

W. Oelert, arXiv: 1501.05728 (2015).

ELENA Upgrade – 100 keV antiprotons



Separation of the AD bunch into 4 bunches (space charge limit)

Supply of 4 experiments simultaneously 24/7 beam for four experiments

Supply of the GBAR experiment from next year

All other experiments from 2021

Next step – collect all the antiprotons!



Important Question: What do we do with all the antimatter?

Of course...



EXTREME PHYSIK

In Dan Browns Bestseller villaminatik planen Verschwören, den Vatikan auszulöschen. Dass sie dallär Antimaterie verwenden wollen, mag als reihe Sciencelliction erscheinen. Aber der Ehdruck täuscht: Die Forscher entdecken Immer neue Einsatzmöglichvon abszei**WIR JAGEN DEN VATI-**

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How much antimatter do we get?

30 million antiprotons in 2 minutes



 8×10^{12} antiproton per year

 $E_{annihilation} = 2mc^2 = 2.4 \text{ kJ/year}$

Your antimatter bomb can...

...prepare one expresso about every leap year







25 ml of water / 88 degrees

How to store (anti)particles?



A plasma of charged (anti-)particles can be confined in a Penning-Malmberg trap

Radial confinement by magnetic field

Electrostatic confinement by axial electric field

Extremely good vacuum (p << 10⁻¹⁰ mbar)

Antiprotons with 5.3 MeV are still to fast to be confined by the electric potential (< 10 kV)

Deceleration by "degrader foils"



How thick do you make the foil to maximize the antiproton yield?

Antiproton scattering



Best choice:

Exit surface is placed in the plane of the mean stopping range

Concerns:

50 % losses by annihilation in the degrader

Broad energy distribution

Beam divergence increases

Bethe-Bloch Formula

Energy loss due to scattering of the projectile with electrons in the material



$$-\frac{dE}{dx} = \frac{\kappa}{\beta^2} n_e Z_1^2 \cdot L(\beta)$$

$$\kappa = \frac{e^4 c^2}{4\pi \varepsilon_0^2 m_e} \quad \text{constant from scattering theory}$$

stopping number

Electron treated as "quasi-free" particle at rest

Projectile energy >> electron binding energy Projectile velocity >> electron velocity

The zero-order term

$$L_0(\beta) = Ln \left[\frac{2m_e c^2}{\langle I \rangle} \frac{\beta^2}{(1-\beta^2)} \right] - \beta^2 - \frac{C(\beta)}{Z_2} - \frac{\delta(\beta)}{2}$$

Empirical or model based data is required for:

- The average excitation/ionization potential: $\langle I \rangle$ Correction for quantized energy of the electrons
- Shell correction $C(\beta)/Z_2$ to account for electron velocity
- Density corrections accounting for polarization of the medium $\left. \delta(eta)/2
 ight.$

The stopping number

The stopping power is usually expanded in a semi-empirical approach to include additional corrections compared to pure Rutherford scattering.

$$L(\beta) = L_0 + Z_1 L_1 + Z_1^2 L_2 + \dots$$

The zero-order term is the scattering term in the Bethe-Bloch formula

The **first-order** term is the Barkas term accounts for the difference in **polarization** of the medium depending on the projectile's charge. **Difference in stopping power for protons and antiprotons!**

The **second-order** term is the Bloch term, which accounts for e.g. **three-particle interactions**

Contributions to the energy loss



J.F. Ziegler, J. Appl. Phys. 85, 1249-1272 (1999). (Author of SRIM)

Experimental solution #1

Modulation of stopping power by pressure changes in gas cells



G. Gabrielse et al., Phys. Rev. A 40, 481 (1989).

Experimental solution #2



5.3 MeV



< 10 keV



Degrader with variable thickness on a linear stage

Experimental solution #3

- A "mesh degrader"
- Stack of thin meshes tilted against each other



 generates a "quasi-random" pattern with a structure small compared to the beam diameter

An antiproton catching trap



S. Sellner et al., New J. Phys. 19, 083023 (2017).

Feedthroughs

How to get vacuum $p < 10^{-17}$ mbar?

- Pinch-off technique
 - Pump the trap chamber to $p < 10^{-6}$ mbar
 - Hermetically seal the vacuum chamber by "pinch off" of the pump connection
- Cool down to 4 K
 - No hydrogen diffusion in the walls
 - Freeze out of all gases except helium
 - Even helium attaches in a monolayer to surfaces by van-der-Waals force (T < 20 K)
- Pressure limits set by antiproton annihilation rates to p $\sim 10^{-18}$ mbar

Antiproton window (25 um stainless steel) t. Pinch-off connection

How it looks in practice?

Indium wire as gaskets

Pinch off flange is exchangeable

Tube is hard-soldered into the flange

3 days pumping

pinch-off

experiment installation (1 day)

Cooling to 4.2 K - 24 h



The pinch-off





After installation





Reservoir trap

Antiprotons stored from 03.11.2015 – 22.12.2016





- Storage of antiprotons for more than one year: **405.5 days**
- Extraction of single particles by a potential tweezer scheme

C. Smorra et al., Int. J. Mass Spectr. 389, 10 (2015).S. Sellner et al., New J. Phys. 19, 083023 (2017).

Inversion of the baryon asymmetry: Antibaryon density: ~ $10^8/cm^3$ V < $(50 \ \mu m)^3$ Baryon density: ~ $1 \ / \ cm^3$ p < 10^{-16} Pa

Sympathetic electron cooling

- Cooling from 1 keV to 1 10 meV (10 – 100 K) by sympathetic cooling with electrons
- Electron synchrotron radiation: Cooling time constant ~0.2 s in 1 T



C. Smorra et al., Eur. Phys. J ST 224, 3055-3108 (2015).

Sympathetic cooling limits

Coulomb interaction reduces due to centrifugal separation

Electron equilibrium temperature can be higher that the environment temperature

Radiofrequency noise can excite plasmas in a broad frequency range

Electron kickout

- Elevate the trap potential
- Open for the potential well for a short time (~ 100 ns)
- Electrons are released from the trap
- Antiprotons (and negative ions) remain



Resistive Cooling

the image-current induced by the antiproton motion $P = R_p I^2$ 4 K 300 K Rp AMP AMP FFT Limits: Cooling time: 10 ms (best axial) 20 mins (bad cyclotron) Temperature: 5 – 10 K

A resistance in parallel to the trap electrodes damps

H. Nagahama et al., Rev. Sci. Instr. 87, 113305 (2016).

The challenge: Cooling by ~10 orders of magnitude

Antiproton production: ~ 3.7 GeV

AD final energy: ~ 5.3 MeV

Trap injection energy: < 5 keV

Present cooling limit: 0.5 meV

rf-deceleration

Stochastic Cooling

Electron Cooling

Degrader

Sympathetic cooling Resistive cooling

Laser-cooled ions?

Mission 1 accomplished

- Cold antiprotons are prepared!
 - High-precision measurements on antiprotons
 - First ingredient for antihydrogen production

What about positrons?

Positrons production

• Nuclear beta decay provides a natural source of positrons

$$(Z,A) \rightarrow (Z-1,A) + e^+ + \overline{\nu}_e$$

- Three-body decay with about 1 MeV energy release
- Production energy and energy spread unfavourable for trapping



Isotope	Q-value (MeV)	Half-life
²² Na	0.54	2.6 yr
⁵⁸ Co	0.47	70.8 d
⁶⁴ Cu	0.65	12.7 h
¹¹ C	0.96	20.4 min

Positron interaction in solids

- Positron diffuse in the solid
- Energy loss by inelastic scattering



- At low energies:
- Direct annihilation with conduction band electrons in metals (τ < 1 ps)
- Positronium formation and annihilation (τ = 142 ns or τ = 124 ps)
- Trapping of positions in crystal defects and annihilation

We need a non-conducting, defect-free crystal with high threshold for positronium formation

Moderation of positrons

Moderators to Produce Slow Positrons

Frozen noble gases (Ne, Ar, Kr) most efficient



Metals (Cu, W) are reasonably good moderators too but ≤ 0.1 % efficient

- Solid noble gases have a large band gap
- Positrons can only thermalize by phonon emission
- Annihilation and positronium formation surpressed



Courtesy of C. Surko

A.P. Mills et al., Appl. Phys. Lett. 49, 1121 (1986).

Energy spectrum after moderation



Positron energy: ~ 1 eV

Energy spread: ~ 0.5 eV ~ 6000 K

Surko-type positron accumulator



Typical temperatures in an accumulator: 10 K – 100 K

http://alpha.web.cern.ch

Some notes on positronium

- Positronium is a hydrogen-like atom
- Spectroscopy on positronium allows high-precision tests of quantum electrodynamics
 - No uncertainties due to finite proton size!
- Most accurate determination of the positron mass comes from the 1S-2S spectroscopy of positronium
 - High-precision CPT test with leptons ($\delta m/m \approx 8 \ 10^{-9}$)
- Rydberg positronium can be used to produce cold antihydrogen by charge exchange:

$$Ps^* + \bar{p} \Rightarrow e^+ + \bar{H}^*$$

e-

e+

Mission 2 accomplished

Cold positrons can be prepared! Now we can make antihydrogen!

TO BE CONTINUED

Antihydrogen Experiments